

Week 4 Discussion

Friday, April 23, 2021 3:53 PM

Time / Age

- lookback time

* cosmic expansion $\Rightarrow H(t), a(t), z(t)$

$\Rightarrow H(a), H(z)$

$$H(t) = \frac{\dot{a}}{a} = \frac{da}{a dt} \Rightarrow dt = \frac{da}{H a}$$

$$\begin{aligned} \int_{t_e}^{t_0} dt &= \int_{a_e}^{a_0} \frac{da}{H a} = \int_z^0 \frac{1+z}{H} d\left(\frac{1}{1+z}\right) \\ &= \int_z^0 \frac{-(1+z)(1+z)^{-2}}{H} dz = \int_0^z \frac{dz}{H(1+z)} \end{aligned}$$

$$\Rightarrow t_0 - t_e = \int_0^z \frac{dz}{\underbrace{H(z)(1+z)}} \rightarrow \text{lookback time}$$

$$\underline{H^2(t) = \frac{8\pi G}{3c^2} \rho(t)}$$

For a matter-dominated universe ($\Omega_m = 1$)

$$\rho \propto \rho_0 (1+z)^3$$

$$\Rightarrow \rho(t) = \frac{\rho_0}{a^3} = \rho_0 (1+z)^3$$

$$H^2(t) = H_0^2 / a^3 = H_0^2 (1+z)^3 \Rightarrow \underline{H(z) = H_0 (1+z)^{3/2}}$$

$$t_0 - t_e = \int_0^z \frac{dz}{H_0 (1+z)^{5/2}} = \frac{2}{3H_0} [1 - (1+z)^{-3/2}]$$

Distances

(1) physical / proper distance

$$ds = c dt = a \underbrace{dr}_{\text{co-moving distance}}$$

(2) luminosity distance

observed flux $l = \frac{L \rightarrow E/t}{4\pi (d_L)^2}$

$E/t \cdot A$ ←

$$\downarrow \textcircled{E} = \frac{hc}{\textcircled{\lambda}} \uparrow$$

① 1

$$= \frac{L}{4\pi (a_0 r)^2 (1+z)^2}$$

- $\frac{1}{1+z}$ (1+z)
 ① $\frac{1}{1+z}$: redshift of photon's λ
 ② $\frac{1}{1+z}$: rate of photon emission

$$\Rightarrow d_L = \underline{a_0 r} (1+z)$$

(3) angular diameter distance



$$s = d_A \cdot \theta = \underline{a(t)} \cdot r \cdot \theta$$

$$\Rightarrow d_A = \frac{a_0 r}{1+z} = \frac{d_L}{(1+z)^2}$$

$$\theta = \frac{s}{r} (1+z) = \frac{s}{a r}$$